Chapter 5 Work, energy, and power

5.1 Energy conservation

Candidates should be able to:

- 1 understand the concept of work, and recall and use work done = force × displacement in the direction of the force
- 2 recall and apply the principle of conservation of energy
- 3 recall and understand that the efficiency of a system is the ratio of useful energy output from the system to the total energy input
- 4 use the concept of efficiency to solve problems
- 5 define power as work done per unit time
- 6 solve problems using P = W/t
- 7 derive P = Fv and use it to solve problems
 - In physics, work (W) is the energy transferred to or from an object through the application of force (F) along a displacement (d)

$$W = F \times d$$

- The SI units for work is in Joules
- The principle of conservation of energy states that energy is neither created nor destroyed. But may transform from one type to another.
- E.g. work can be transformed to heat (friction!), electric to light.
- Not all energy transferred is useful. E.g. when transferring electric to light, some energy is wasted in the form of heat!
- Efficiency is the ratio of useful energy output from the system to the total energy input

$$\eta = \frac{Useful \ Output}{Total \ Input}$$

• Power is the amount of energy (E), transferred per unit time

$$P = \frac{E}{t}$$

- The SI units for work is in Watt or Joules/ second.
- Moving power is defined by the equation

$$P = F \times v$$

- This equation is used when a force is moving a body at a constant velocity
- Below is how you derive the equation

$$P = \frac{E}{t}$$

Substitute E with

$$W = F \times d$$

$$P = \frac{F \times d}{t}$$

Recall d/t is velocity (Chapter 1). Therefore

$$\boldsymbol{P}=\boldsymbol{F}\times\boldsymbol{v}$$

5.2 Gravitational potential energy and kinetic energy



- 1 derive, using W = Fs, the formula $\Delta E_p = mg\Delta h$ for gravitational potential energy changes in a uniform gravitational field
- 2 recall and use the formula $\Delta E_p = mg\Delta h$ for gravitational potential energy changes in a uniform gravitational field
- 3 derive, using the equations of motion, the formula for kinetic energy $E_{\rm K} = \frac{1}{2}mv^2$
- 4 recall and use $E_{\rm K} = \frac{1}{2}mv^2$
 - The change in **Gravitational potential energy (ΔE)** is the energy stored in a **mass (m)** relative to its **position in a gravitational field (Δh)**

$$\Delta E = mg\Delta h$$

Derivation



• Kinetic energy (KE) is the energy an object possesses due to its velocity.

$$KE = \frac{1}{2}mv^2$$

Derivation

Start with $v^2 = u^2 + 2ad$ (Chapter 1) Assume initial velocity is 0 (u = 0) and rearranging $a = \frac{v^2}{2d}$ Recall F = ma (Chapter 2) Substituting $F = m(\frac{v^2}{2d})$ and finally, into the work equation W = Fd We have $W = m(\frac{v^2}{2d})d$ Therefore $W = m(\frac{v^2}{2})$ Or $KE = \frac{1}{2}mv^2$